

# HITS

Heidelberg Institute for  
Theoretical Studies

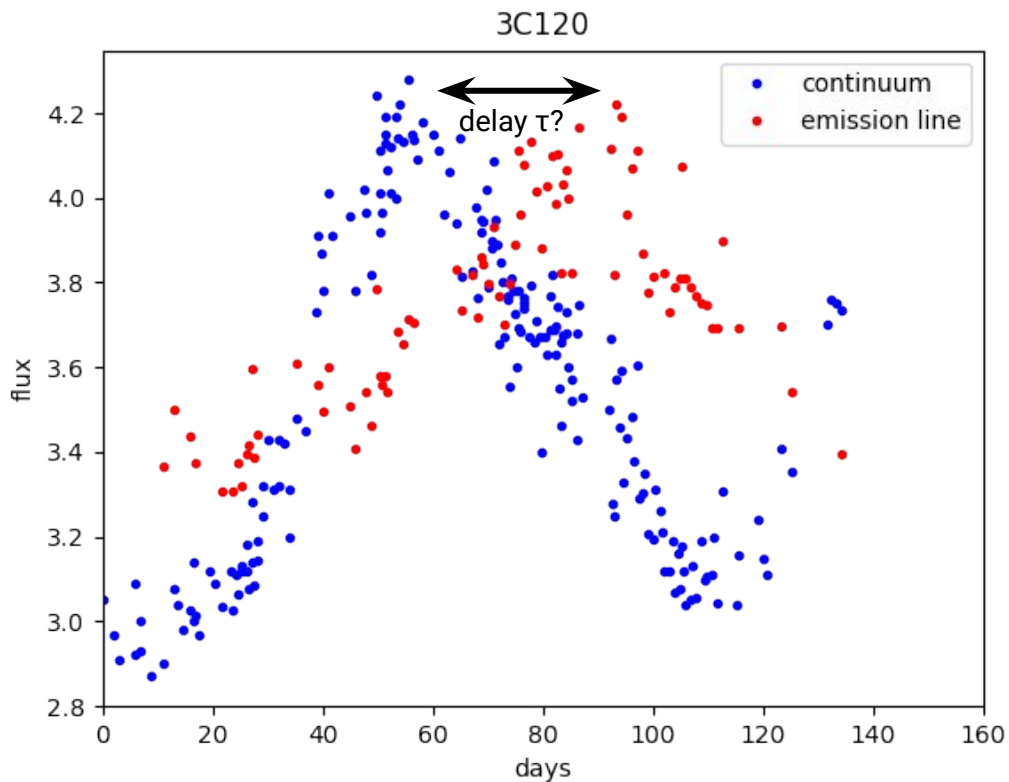
# Probabilistic Cross Correlation for Delay Estimation

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Astroinformatics Group

<https://www.h-its.org/research/ain/>

# Overview: what?



Find the delay  $\tau$  between two lightcurves originating from an AGN

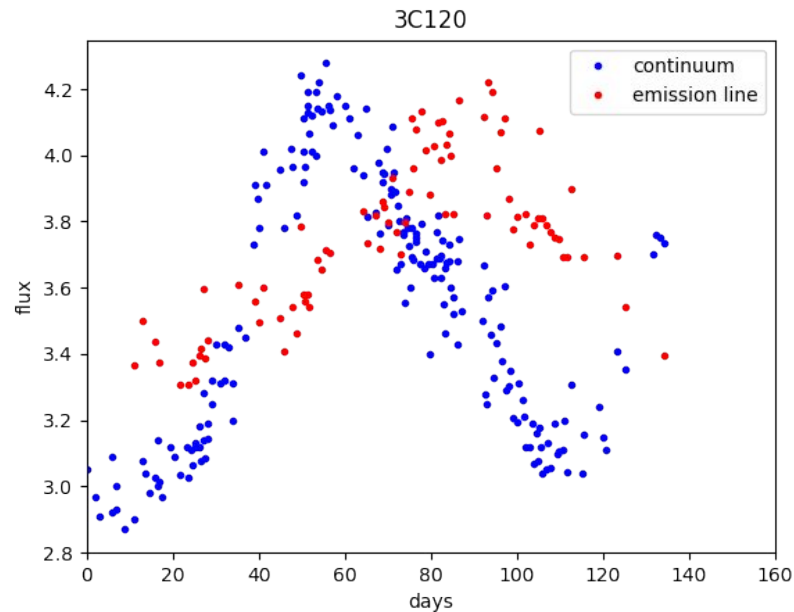
# Overview: why?

- Ultimate goal is to (re)estimate the Hubble constant  
This sheds light on the evolution of the universe
- Estimation of black hole masses in AGN is crucial for this estimation  
Delay  $\tau$  between lightcurves (+ assumptions) helps us infer black hole mass in AGN

# Overview: how?

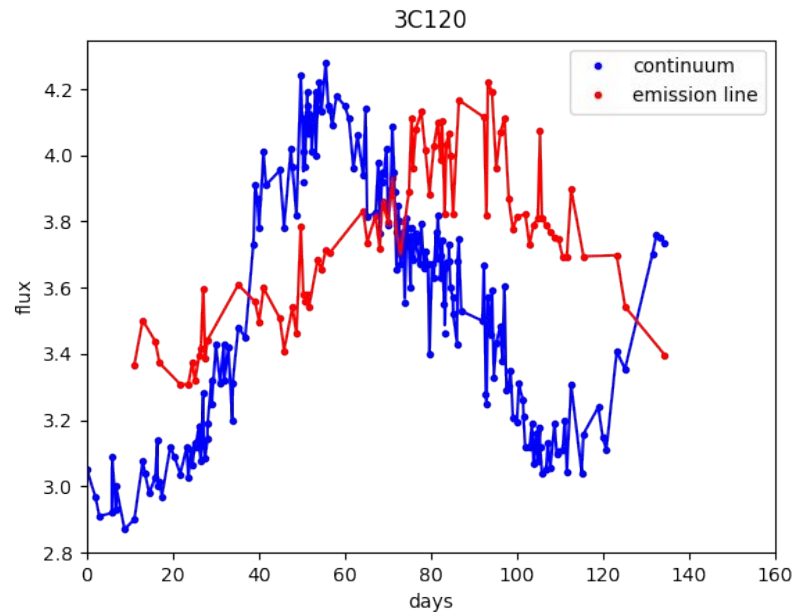
- Interpolated Cross-Correlation Function (and co) estimates delay, but doesn't propagate the uncertainty (measurement noise) present in the data
  - we're uncertain about the data  $\Rightarrow$  delay estimate will be uncertain
- Probabilistic reformulation of ICCF that overcomes this issue
  - uncertainty present in the data is propagated
  - delay estimate described by probability distribution
- We demonstrate our method on object 3C120

# Interpolated Cross Correlation Function



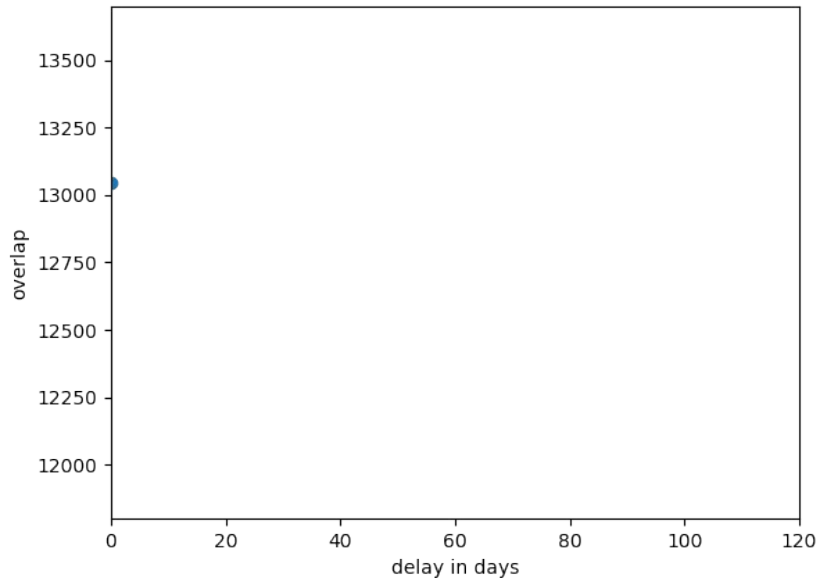
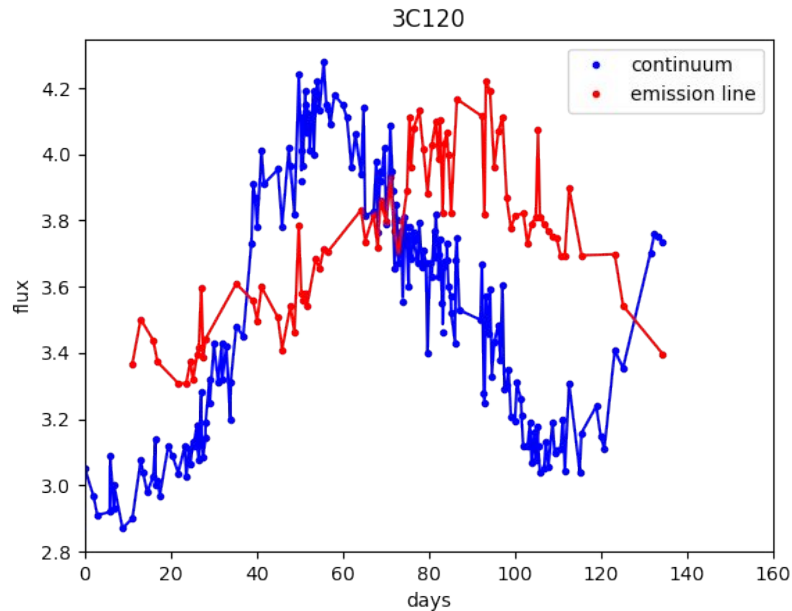
Relies on (linear) interpolation and cross-correlation  $(y_c \star y_e)(\tau) = \int_{-\infty}^{\infty} y_c(t)y_e(t + \tau)dt$   
[Gaskell & Peterson (1987)]

# Interpolated Cross Correlation Function



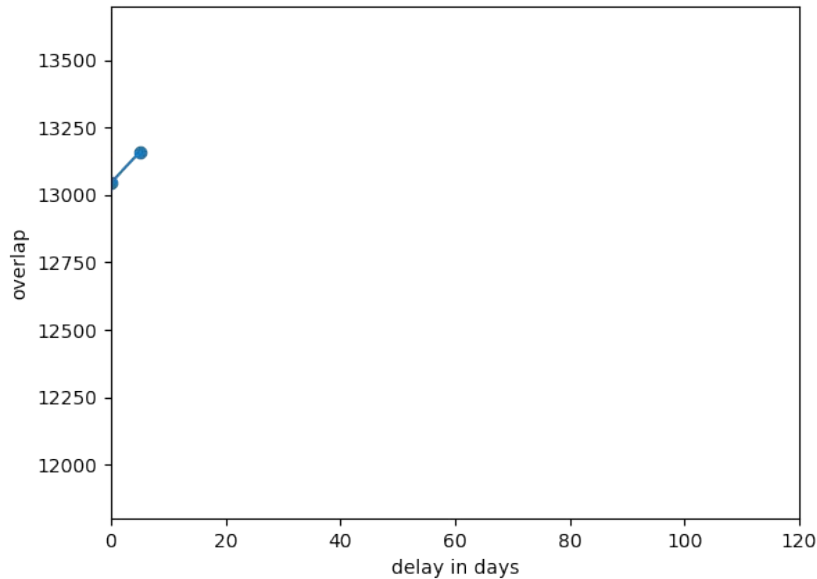
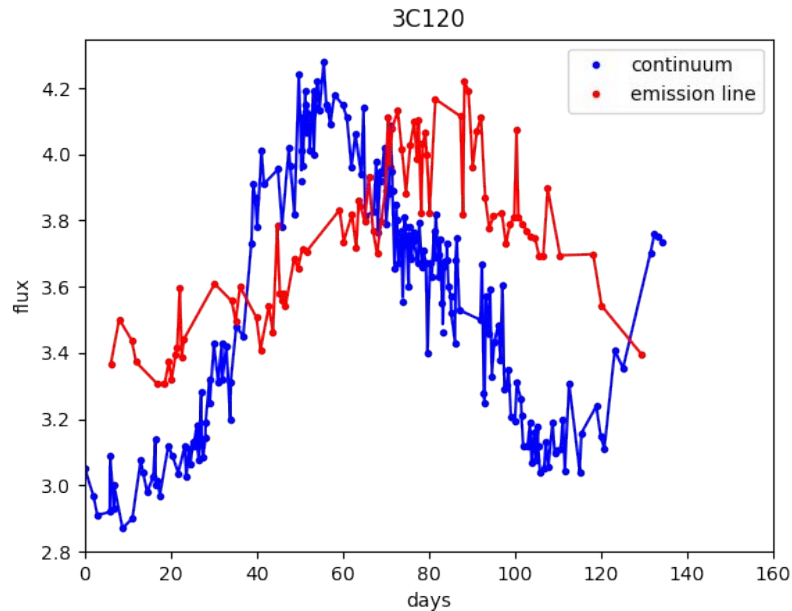
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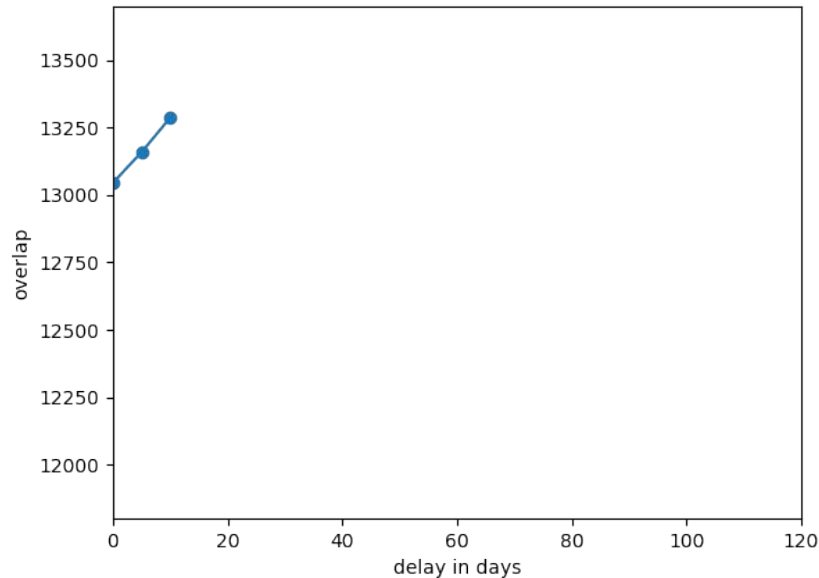
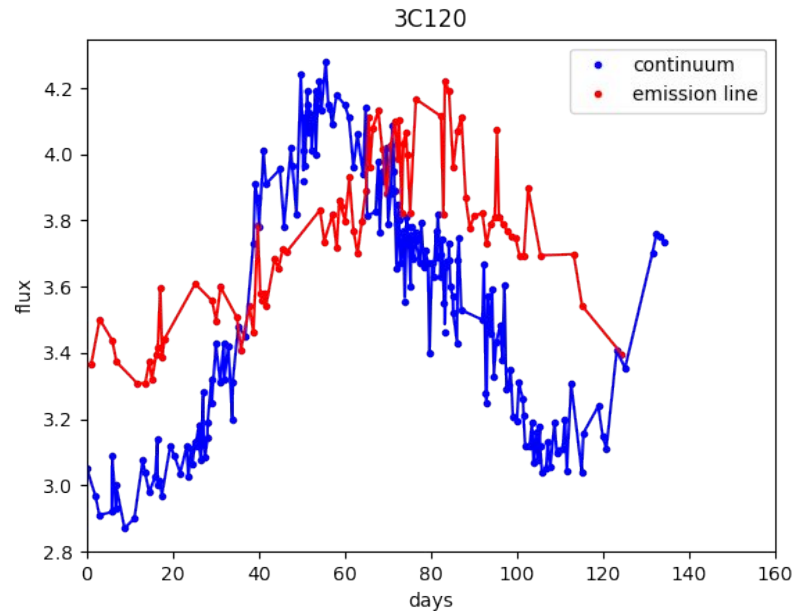
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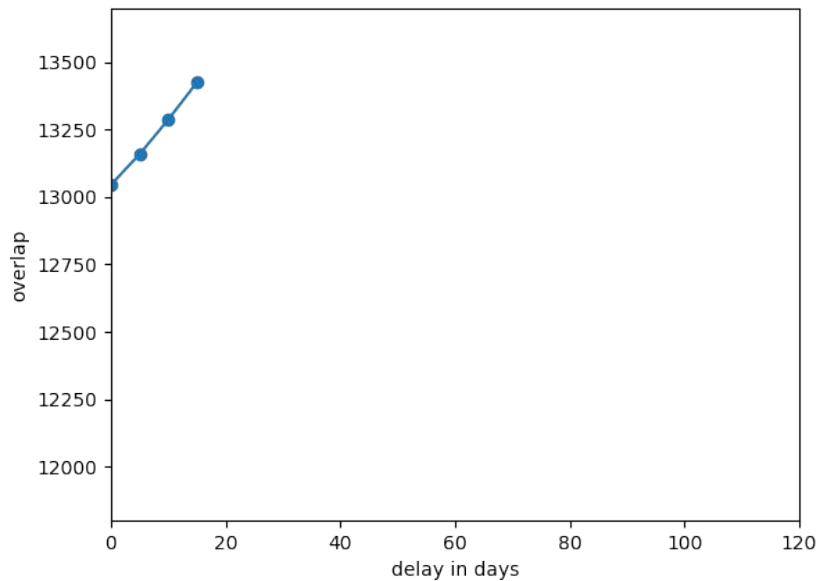
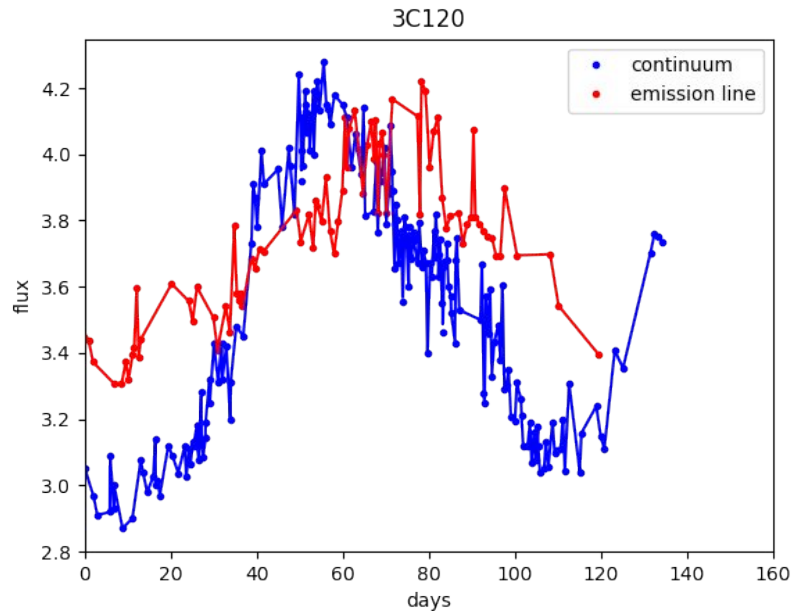


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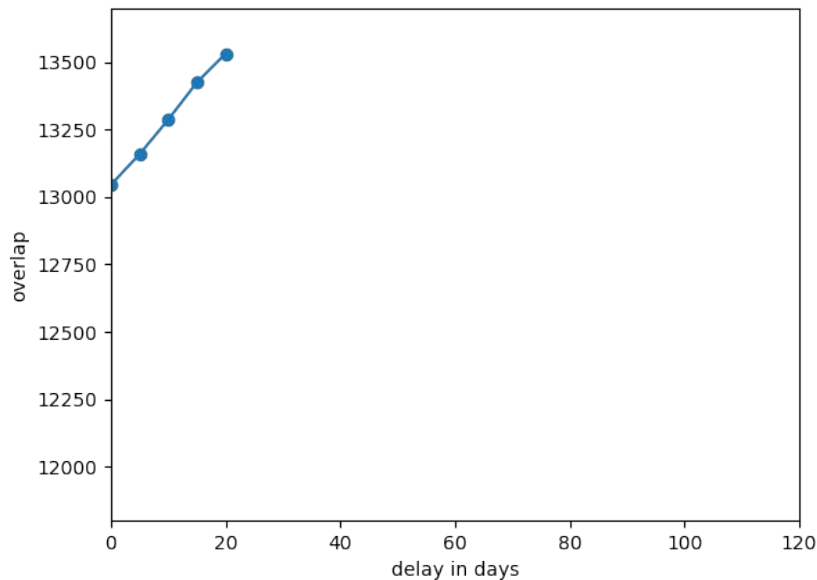
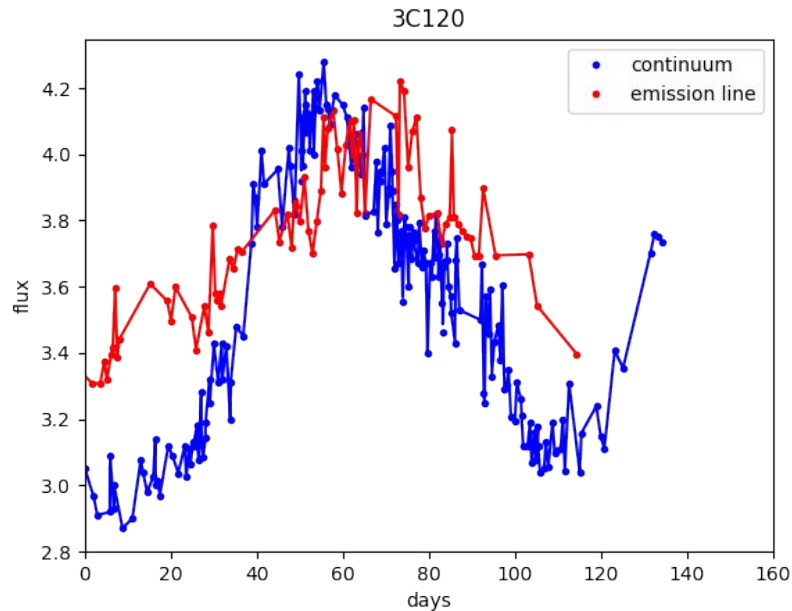
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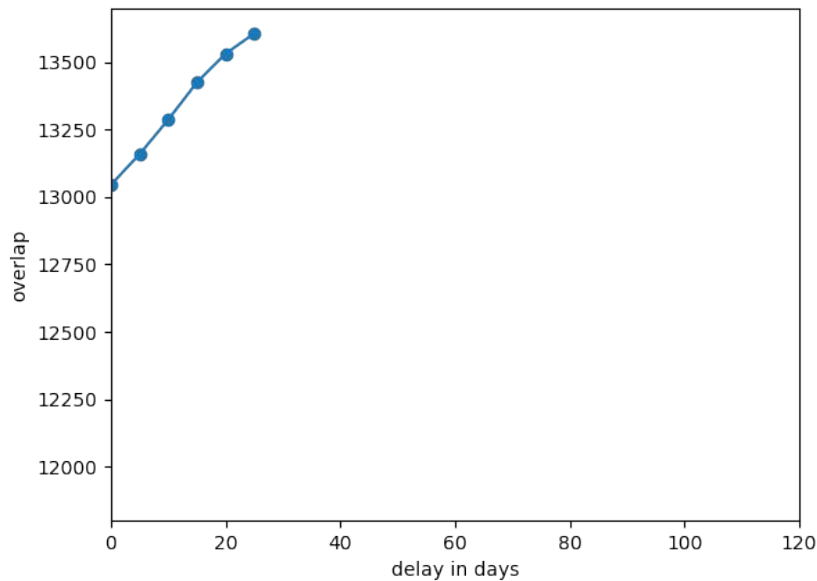
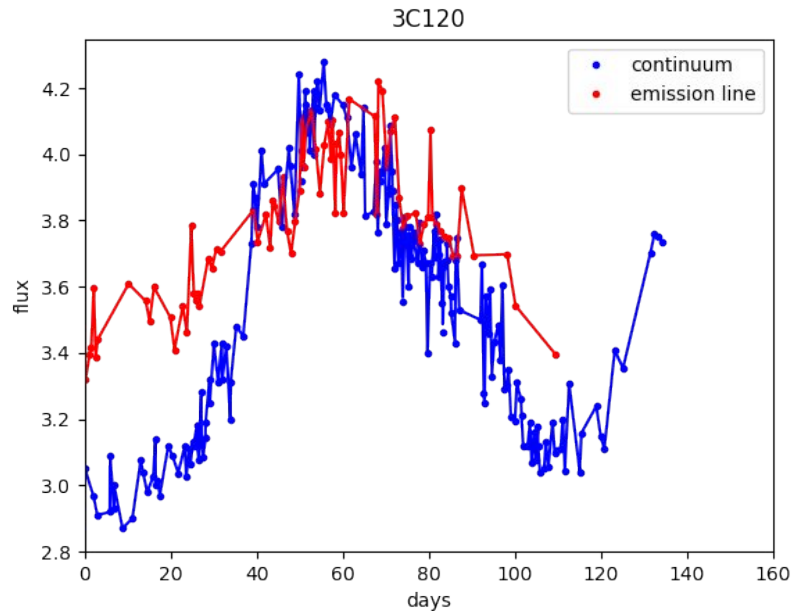
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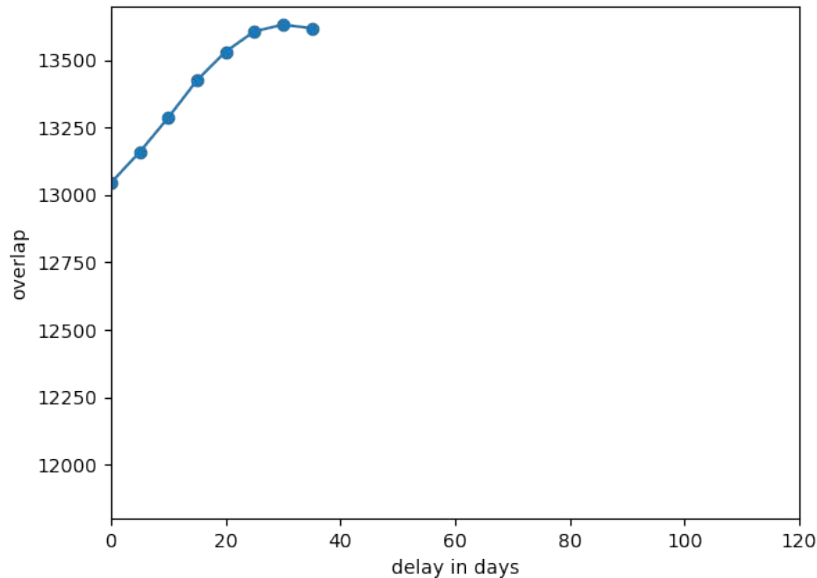
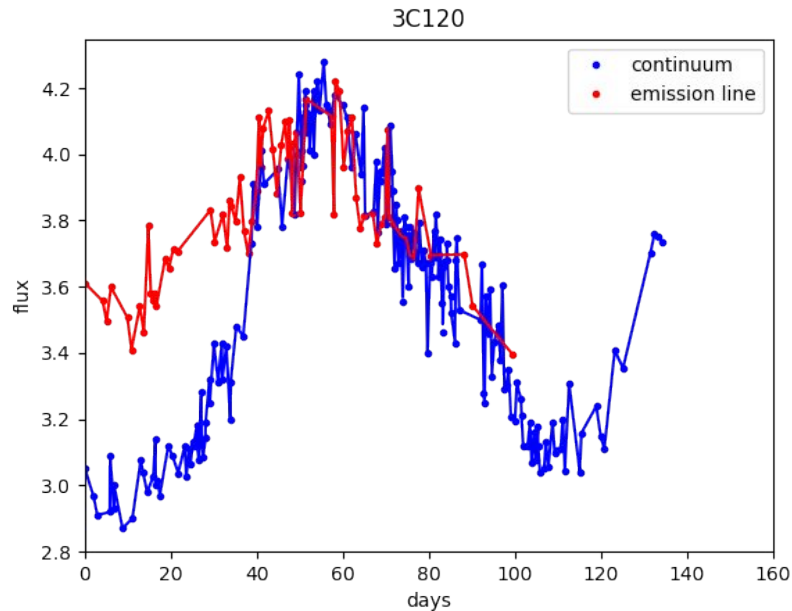
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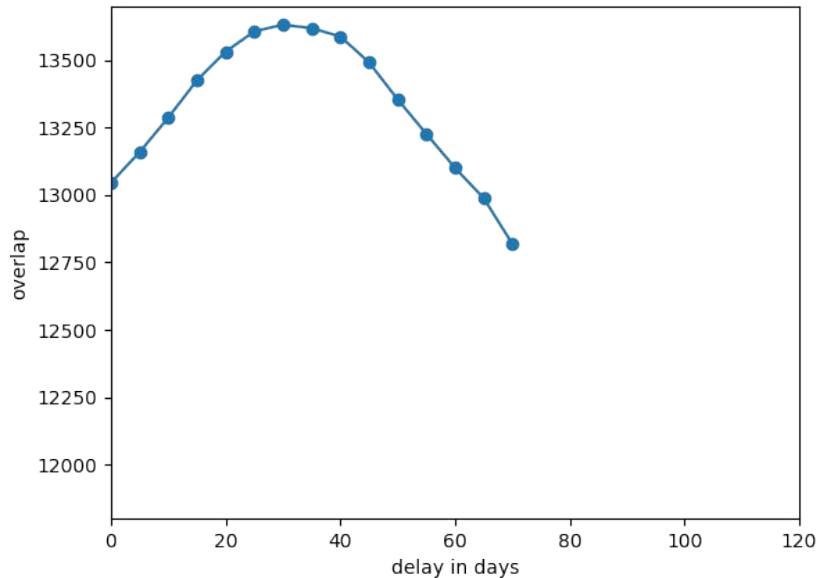
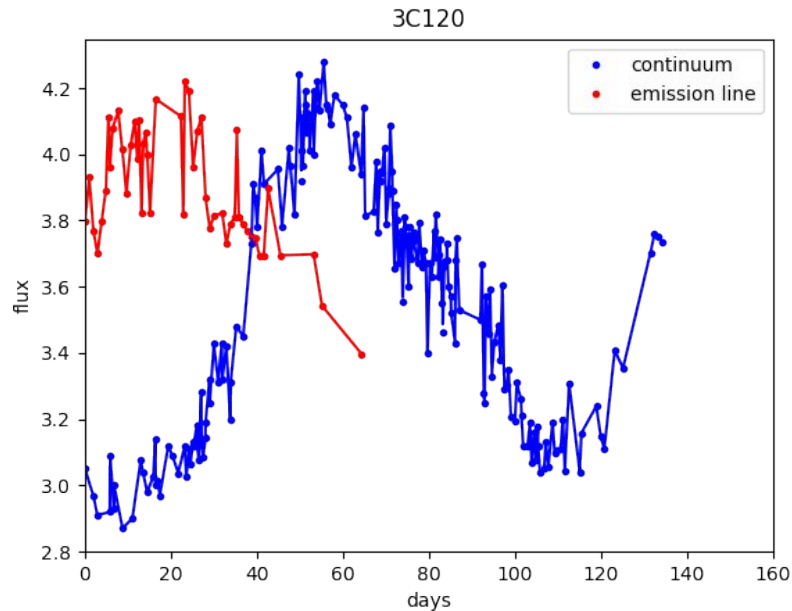
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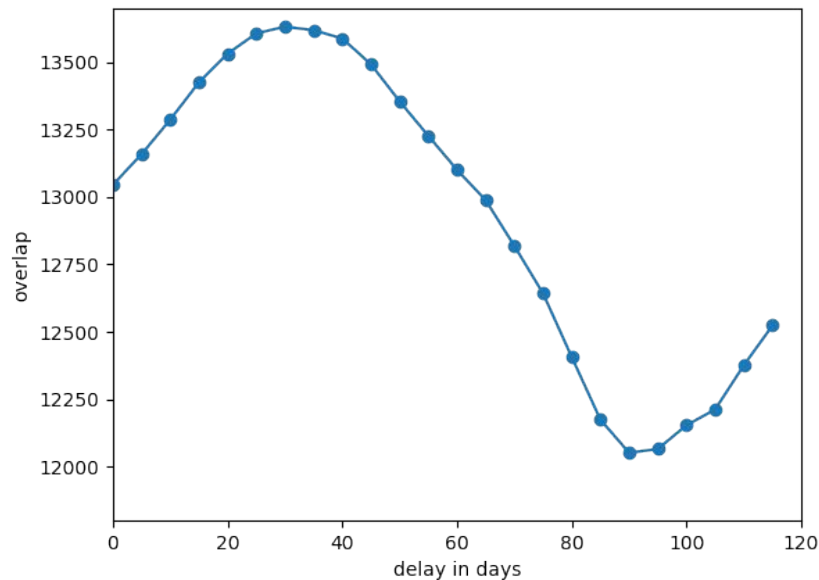
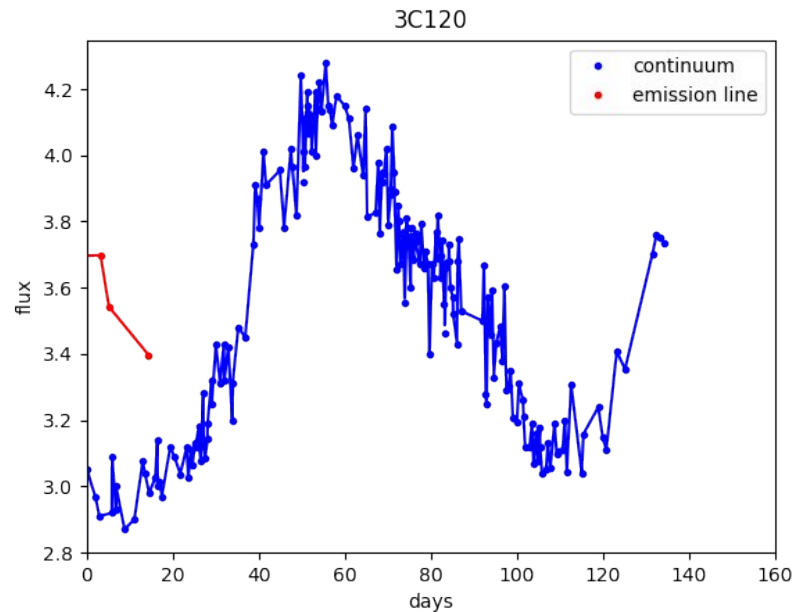
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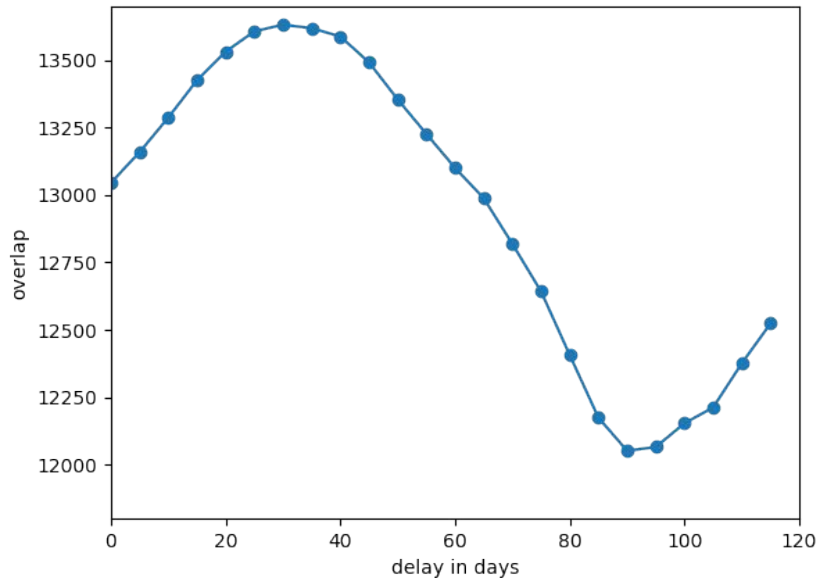
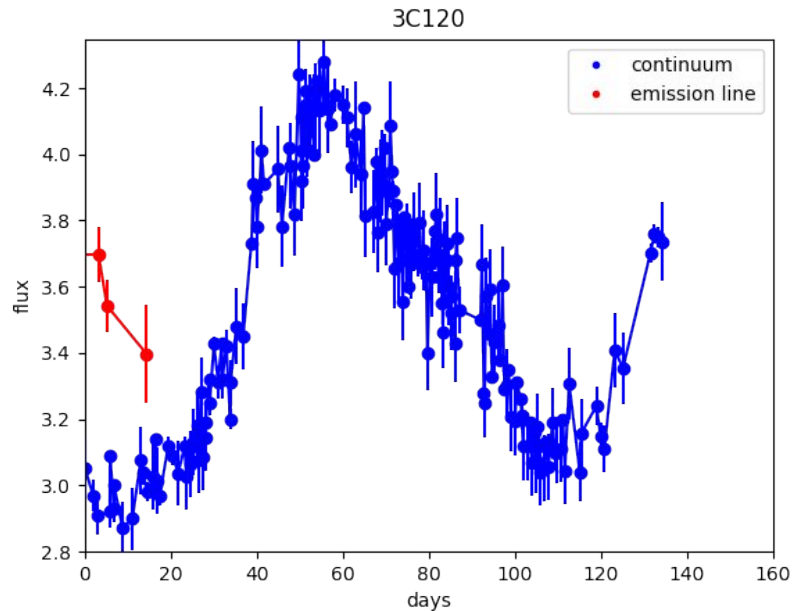
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**ICCF does not propagate noise uncertainty**



# Assumptions behind ICCF

- ICCF finds delay via

$$(y_c \star y_e)(\tau) = \int_{-\infty}^{\infty} y_c(t)y_e(t + \tau)dt$$

- If we accept **cross-correlation** as a means to find the delay, then we assume that one lightcurve is a **delayed**, **scaled** and **offsetted** version of the other

$$y_e(t - \tau) = \alpha y_c(t) + b$$

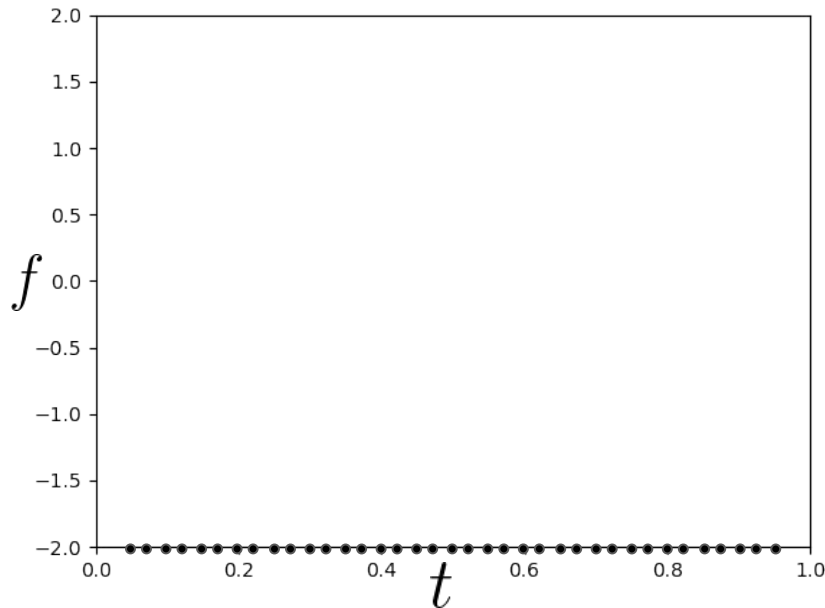
- We can rewrite this relationship in terms of an unobserved, latent signal  $f(t)$

$$y_e(t) = \alpha_e f(t - \tau) + b_e$$

$$y_c(t) = \alpha_c f(t) + b_c$$

# Gaussian Process Modelling

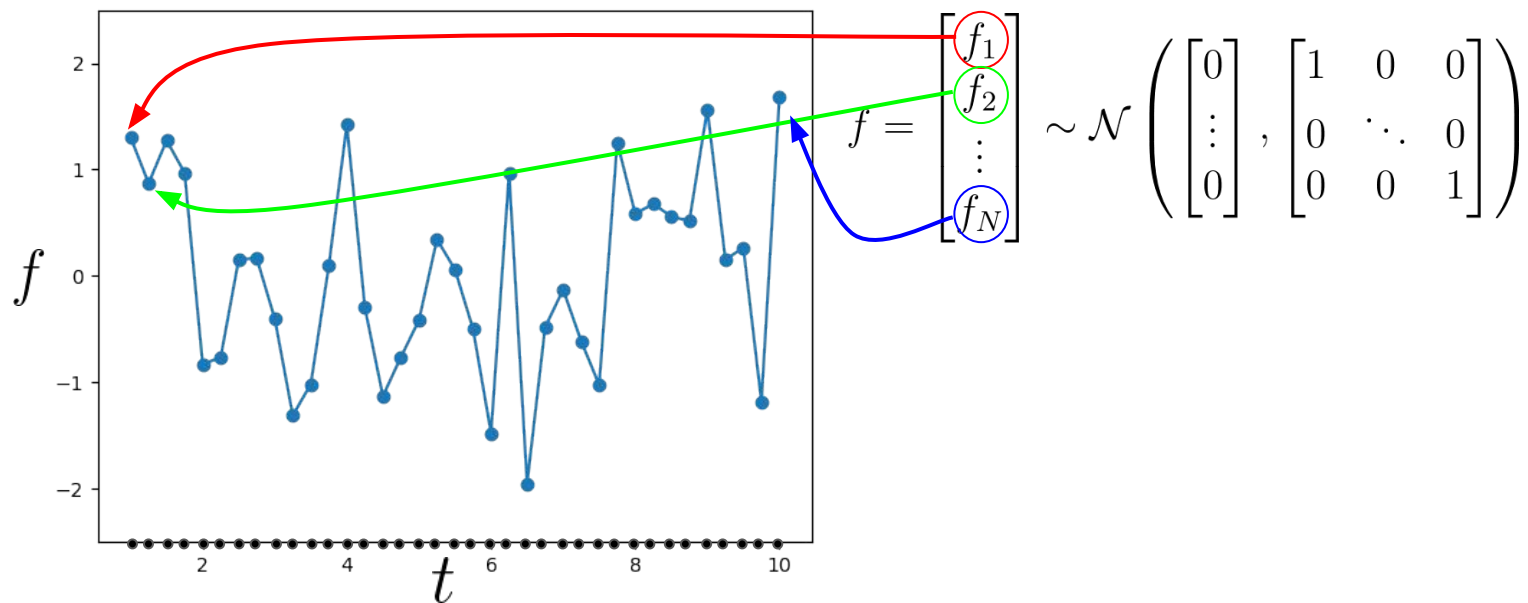
- We model latent signal  $f(t)$  as a sample from a Gaussian process (GP)
- A GP models the values of a function as a Gaussian distribution
- A GP is a tool to model distributions of functions



$$f = \begin{bmatrix} f_1 \\ f_2 \\ \vdots \\ f_N \end{bmatrix} \sim \mathcal{N} \left( \begin{bmatrix} 0 \\ \vdots \\ 0 \end{bmatrix}, \begin{bmatrix} 1 & 0 & 0 \\ 0 & \ddots & 0 \\ 0 & 0 & 1 \end{bmatrix} \right)$$

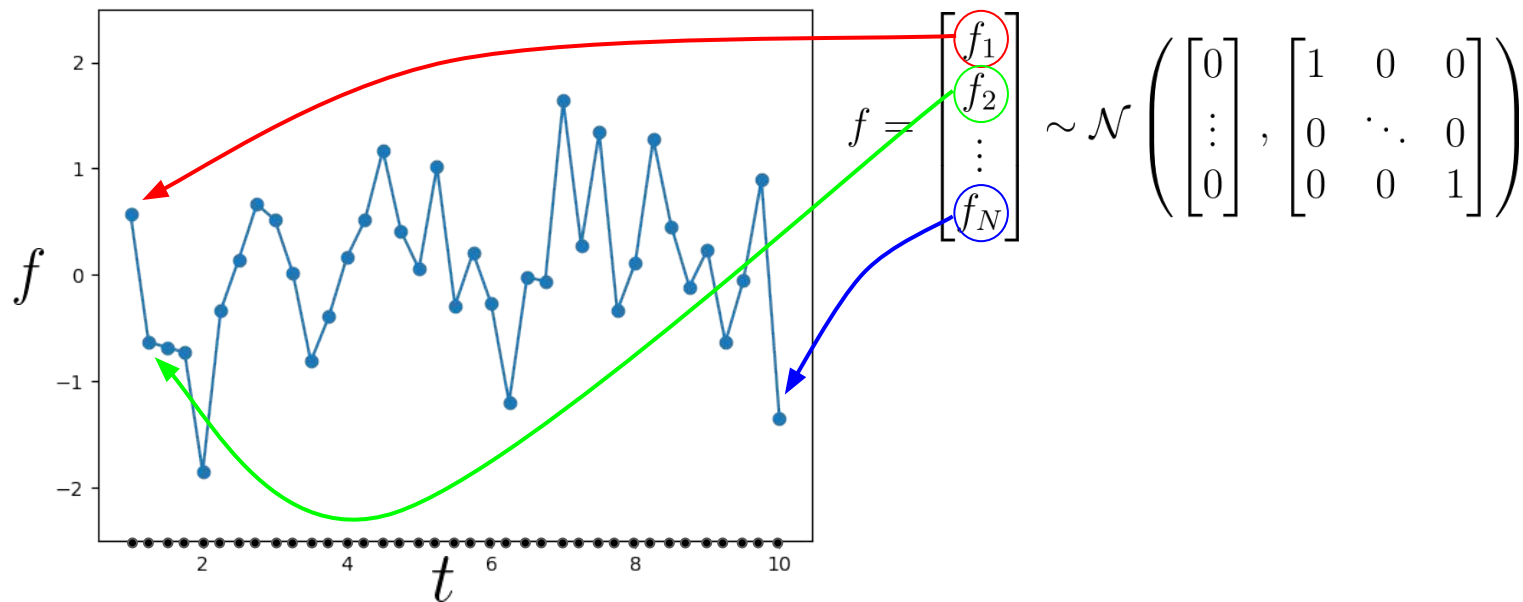
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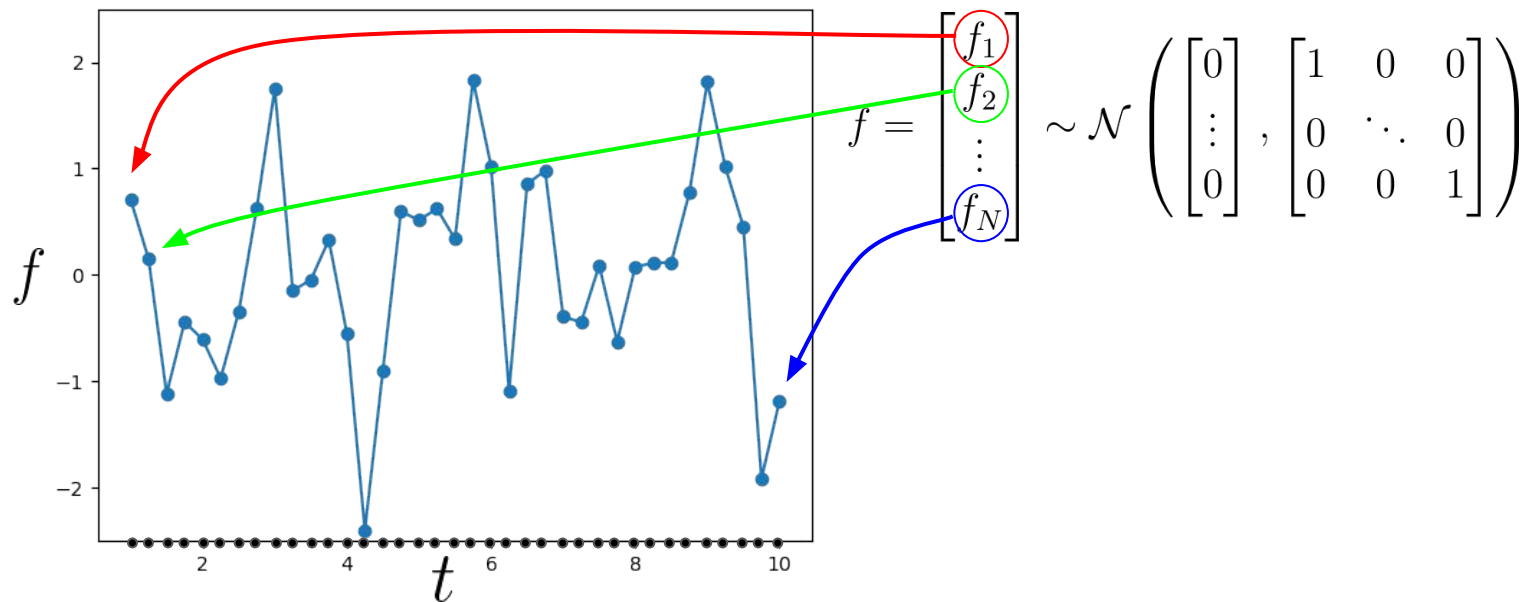
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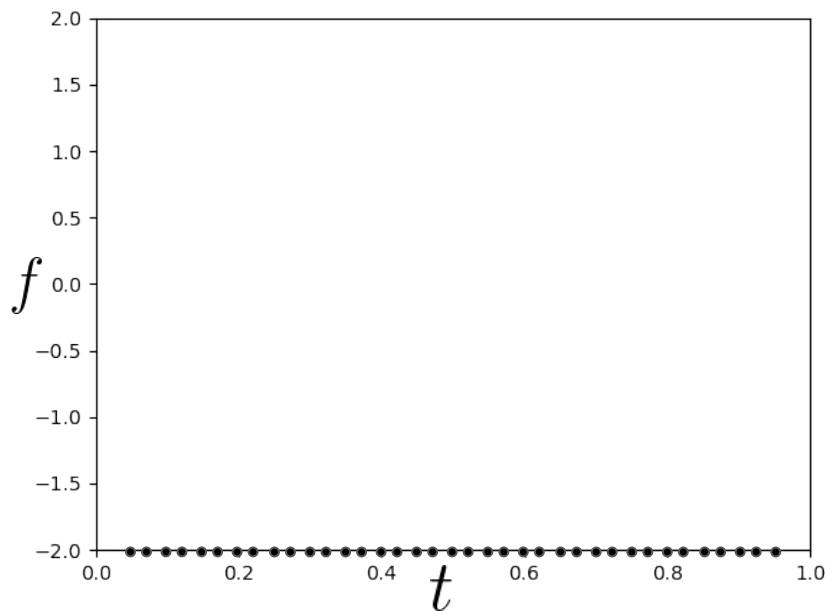
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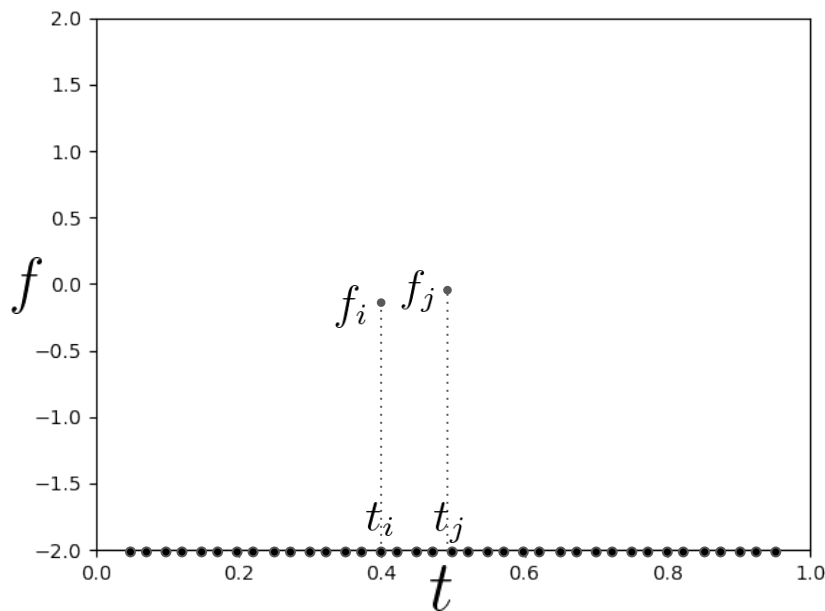


$$f \sim \mathcal{N} \left( \begin{bmatrix} 0 \\ \vdots \\ 0 \end{bmatrix}, \begin{bmatrix} k(t_1, t_1) & k(t_1, t_2) & \dots & k(t_1, t_N) \\ k(t_2, t_1) & k(t_2, t_2) & \dots & k(t_2, t_N) \\ \vdots & \vdots & \dots & \vdots \\ k(t_N, t_1) & k(t_N, t_2) & \dots & k(t_N, t_N) \end{bmatrix} \right)$$

$$k(t_i, t_j) = \exp(-\|t_i - t_j\|^2)$$

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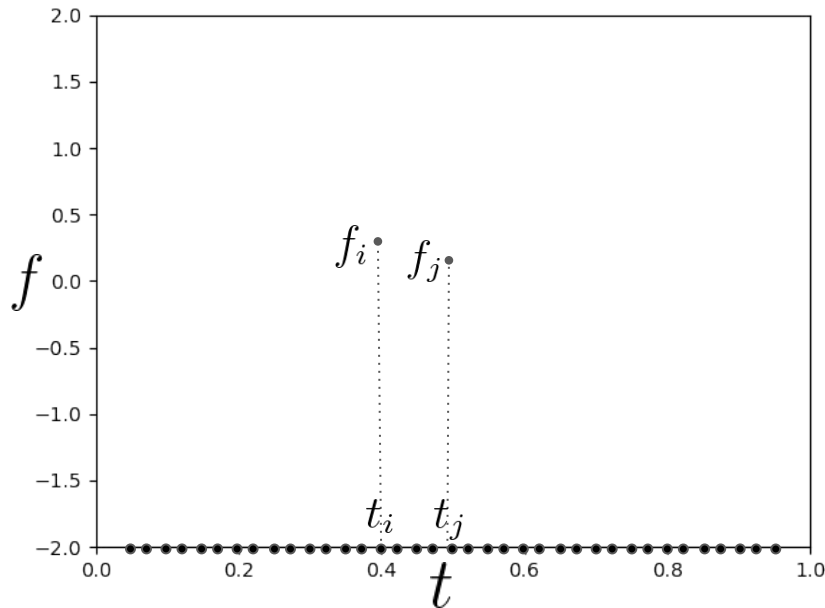


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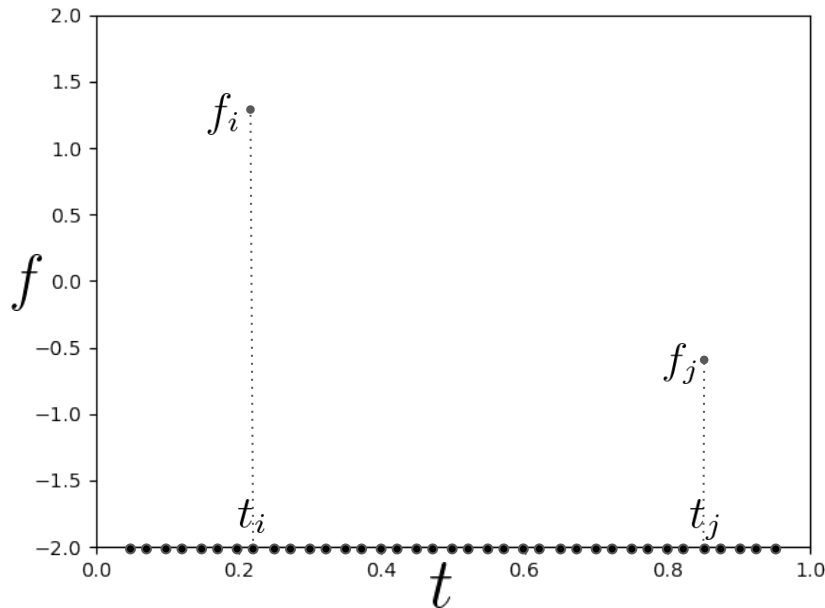
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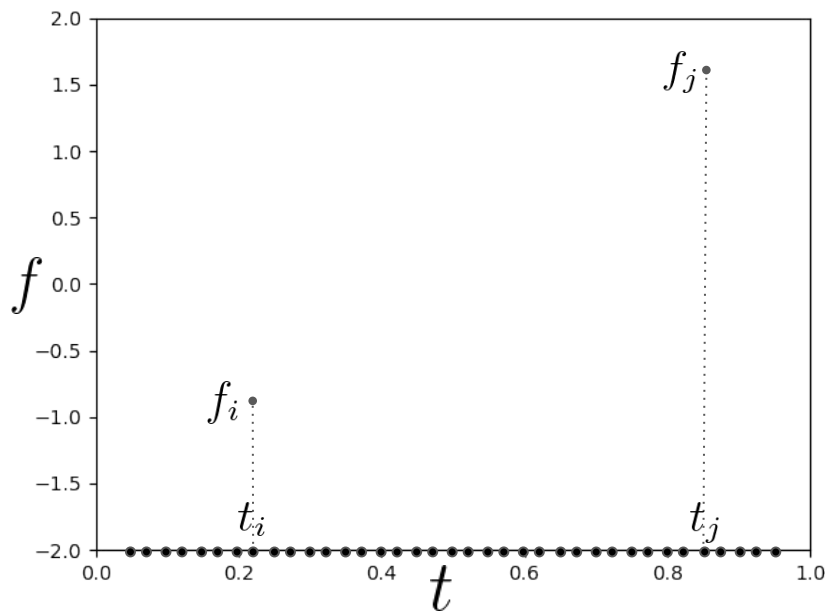


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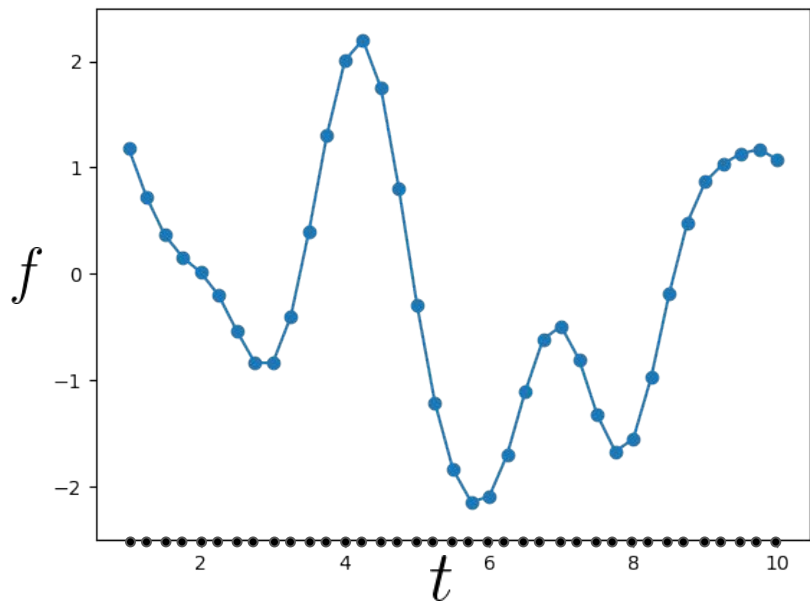


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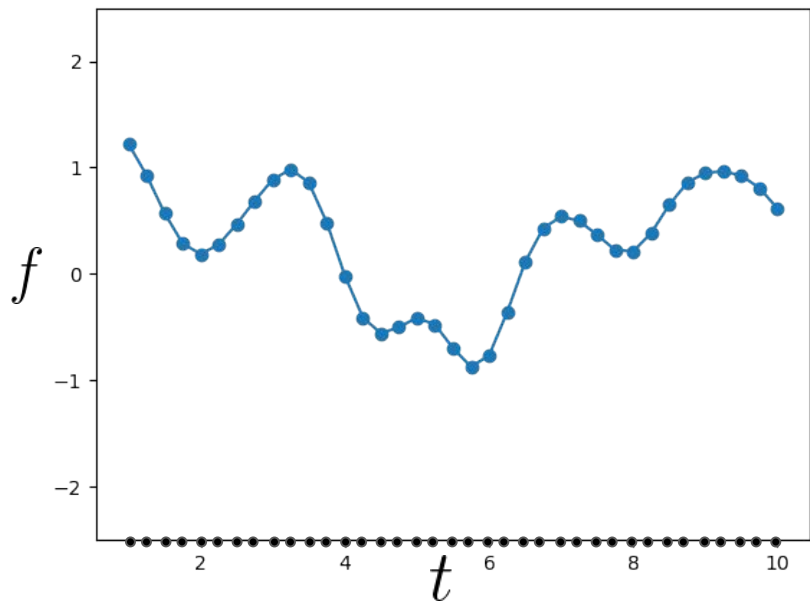


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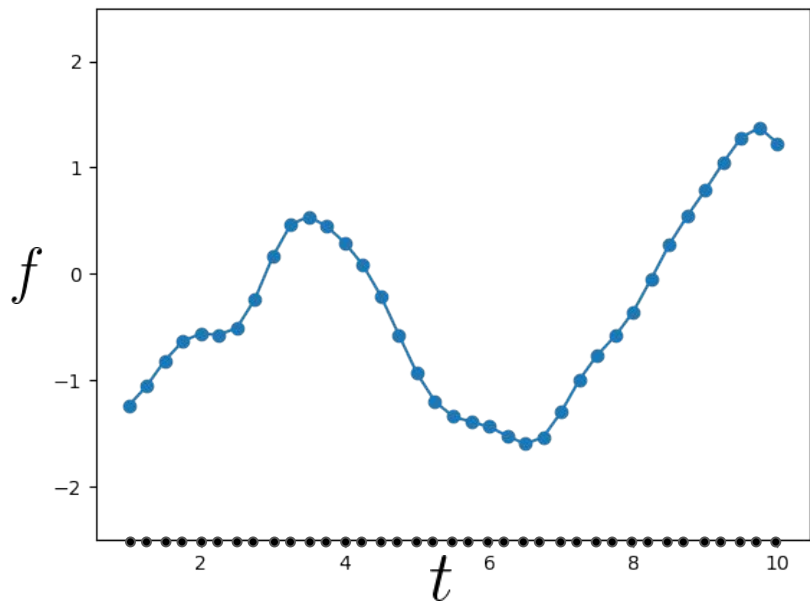


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# Gaussian Process Modelling

- Remember, Gaussian distribution is closed under affine transformations

$$\begin{bmatrix} f_1 \\ f_2 \end{bmatrix} \sim \mathcal{N} \left( \begin{bmatrix} \mu_1 \\ \mu_2 \end{bmatrix}, \begin{bmatrix} \sigma_1^2 & \sigma_{1,2}^2 \\ \sigma_{2,1}^2 & \sigma_2^2 \end{bmatrix} \right) \longrightarrow \begin{bmatrix} \alpha_1 f_1 + b_1 \\ \alpha_2 f_2 + b_2 \end{bmatrix} \sim \mathcal{N} \left( \begin{bmatrix} \alpha_1 \mu_1 + b_1 \\ \alpha_2 \mu_2 + b_2 \end{bmatrix}, \begin{bmatrix} \alpha_1^2 \sigma_1^2 & \alpha_1 \alpha_2 \sigma_{1,2}^2 \\ \alpha_2 \alpha_1 \sigma_{2,1}^2 & \alpha_2^2 \sigma_2^2 \end{bmatrix} \right)$$

- GP also closed under affine transformation

$$f \sim \mathcal{N} \left( \begin{bmatrix} 0 \\ \vdots \\ 0 \end{bmatrix}, \begin{bmatrix} k(t_1, t_1) & k(t_1, t_2) & \dots & k(t_1, t_N) \\ k(t_2, t_1) & k(t_2, t_2) & \dots & k(t_2, t_N) \\ \vdots & \vdots & \dots & \vdots \\ k(t_N, t_1) & k(t_N, t_2) & \dots & k(t_N, t_N) \end{bmatrix} \right)$$

$$\begin{bmatrix} y_c(t_i) \\ \vdots \\ y_e(t_j) \end{bmatrix} = \begin{bmatrix} \alpha_c f(t_i) + b_c \\ \vdots \\ \alpha_e f(t_j - \tau) + b_e \end{bmatrix} \sim \mathcal{N} \left( \begin{bmatrix} b_c \\ \vdots \\ b_e \end{bmatrix}, \begin{bmatrix} \alpha_c^2 k(t_i, t_i) & \dots & \alpha_c \alpha_e k(t_i, t_j - \tau) \\ \vdots & \ddots & \vdots \\ \alpha_e \alpha_c k(t_j - \tau, t_i) & \dots & \alpha_e^2 k(t_j - \tau, t_j - \tau) \end{bmatrix} \right)$$

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- Likelihood reads

$$\begin{bmatrix} y_c(t_i) \\ \vdots \\ y_e(t_j) \end{bmatrix} \sim \mathcal{N} \left( \begin{bmatrix} b_c \\ \vdots \\ b_e \end{bmatrix}, \begin{bmatrix} \alpha_c^2 k(t_i, t_i) & \dots & \alpha_c \alpha_e k(t_i, t_j - \tau) \\ \vdots & \vdots & \vdots \\ \alpha_e \alpha_c k(t_j - \tau, t_i) & \dots & \alpha_e^2 k(t_j - \tau, t_j - \tau) \end{bmatrix} \right)$$

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- or more compact:

$$p(\mathcal{D} | \alpha_c, \alpha_e, b_c, b_e, \tau)$$

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$$p(\mathcal{D} | \alpha_c, \alpha_e, b_c, b_e, \tau)$$

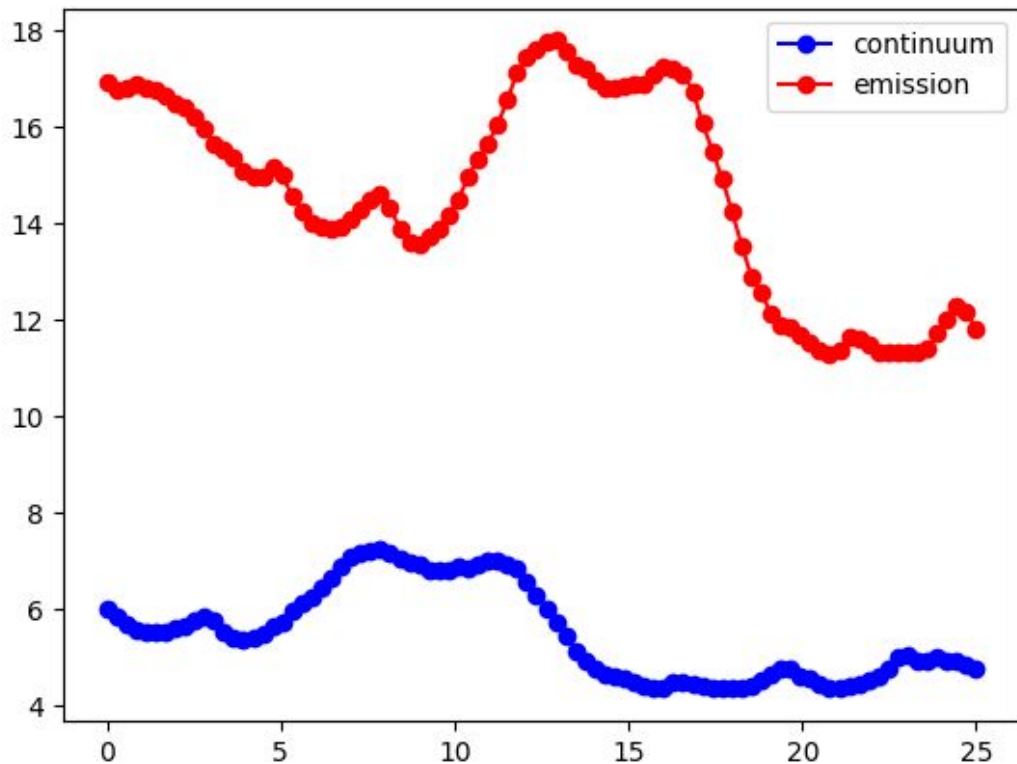
- Posterior delay:

$$p(\tau | \mathcal{D}) \propto \int \overbrace{p(\mathcal{D} | \alpha_c, \alpha_e, b_c, b_e, \tau)}^{\text{likelihood}} \overbrace{p(\tau) p(\alpha_c) p(\alpha_e) p(b_c) p(b_e) p(\theta_\kappa)}^{\text{prior}} d\alpha_c d\alpha_e db_c db_e$$

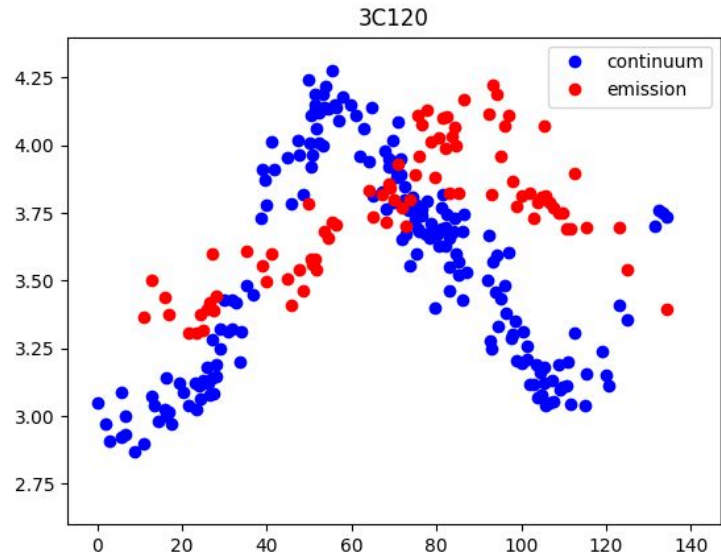
- Integral cannot be analytically done, we resort to approximations not detailed here

# Gaussian Process Cross Correlation

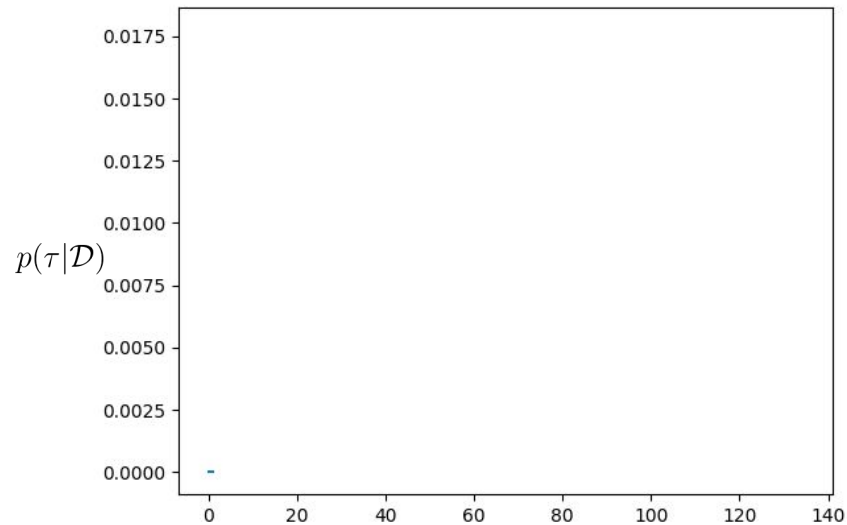
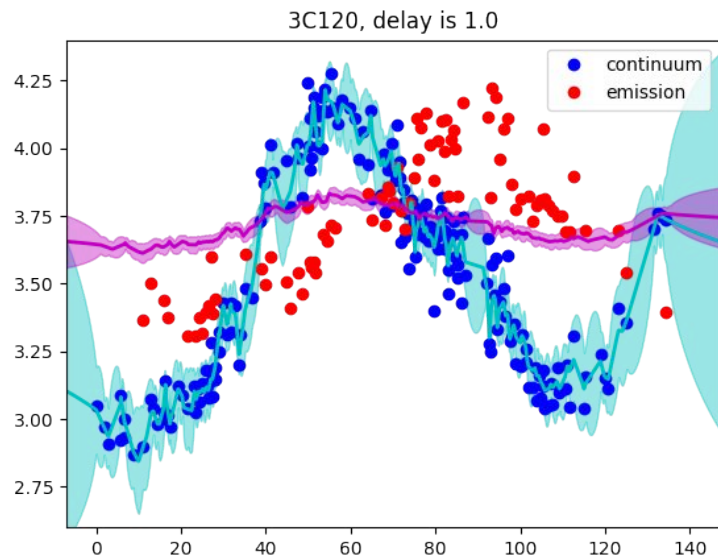
Sampling from our model generates



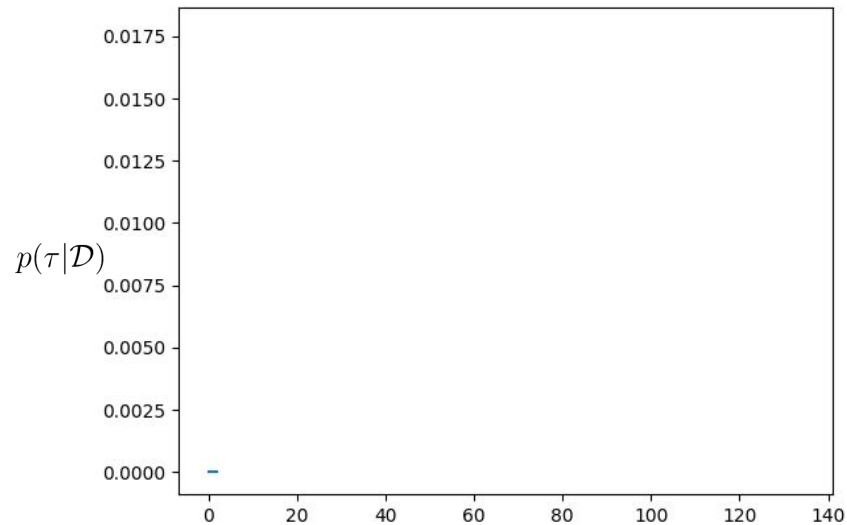
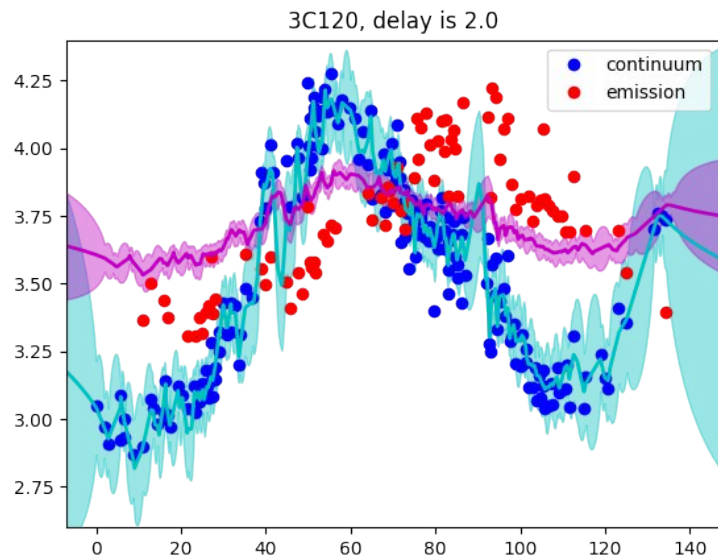
# Posterior delay distribution for 3C120



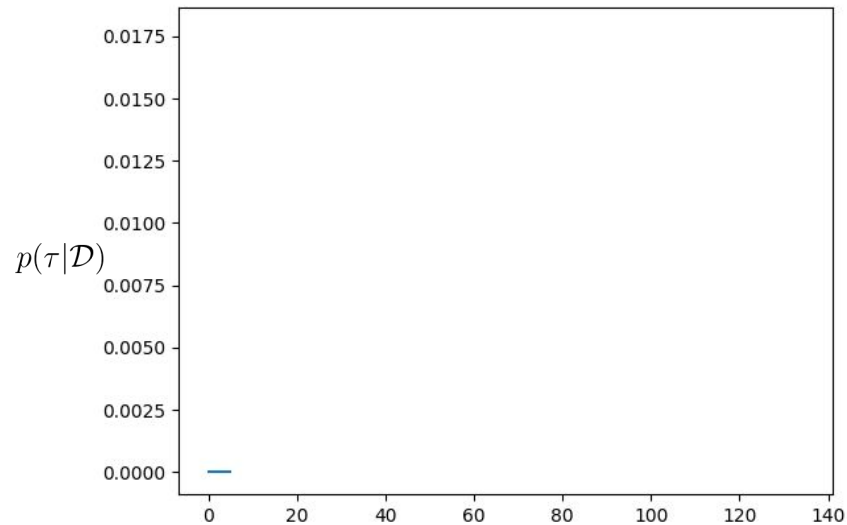
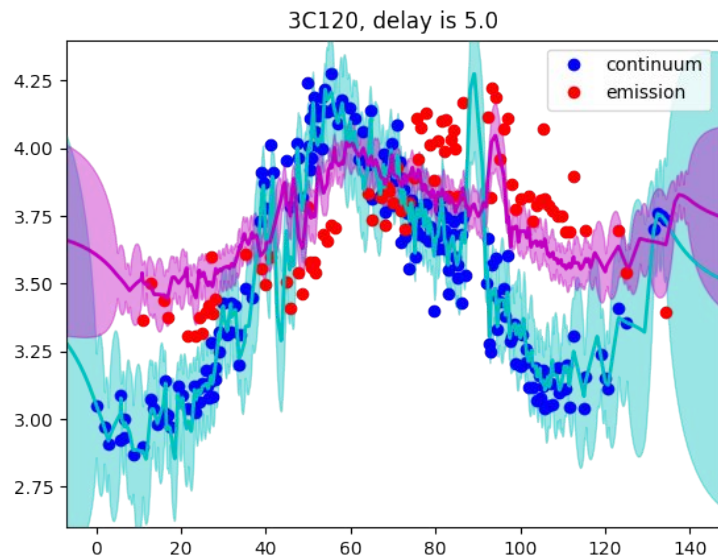
# Posterior delay distribution for 3C120



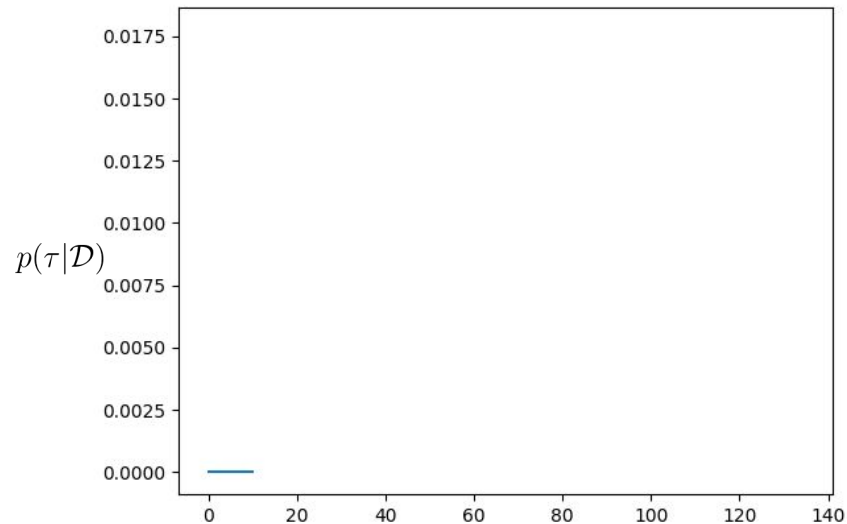
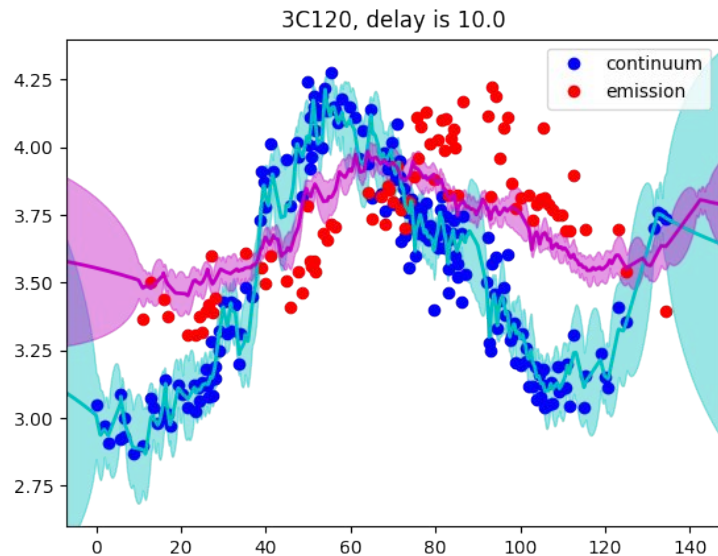
# Posterior delay distribution for 3C120



# Posterior delay distribution for 3C120

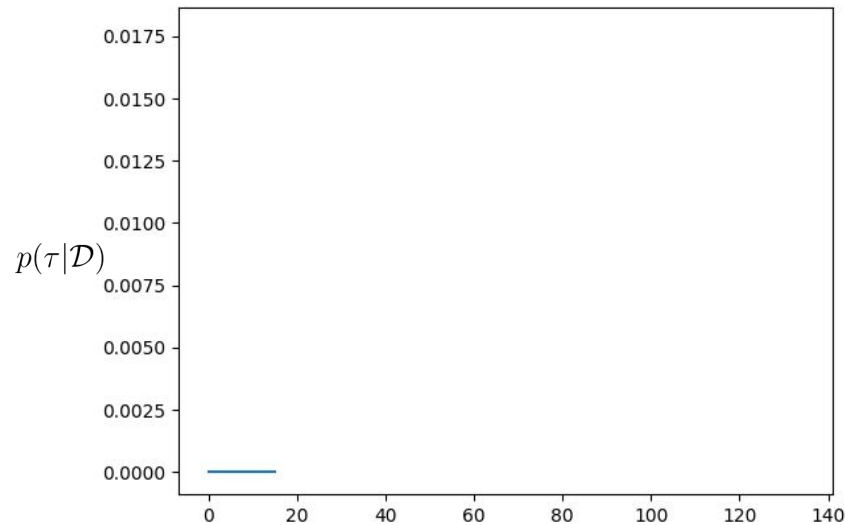
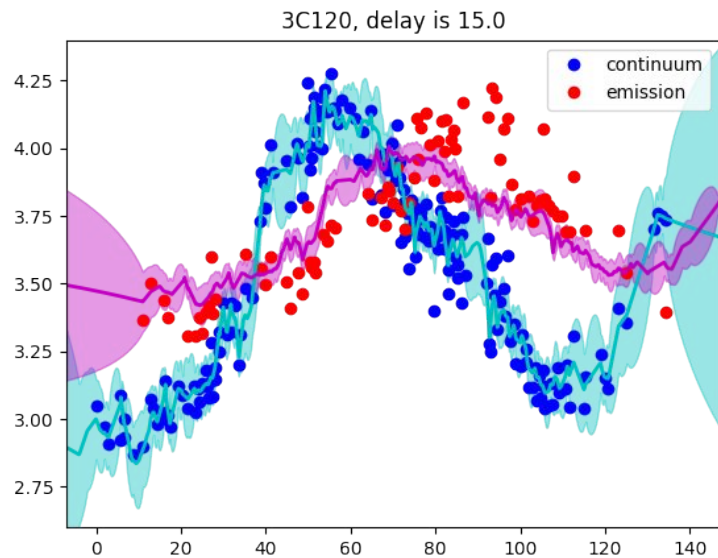


# Posterior delay distribution for 3C120

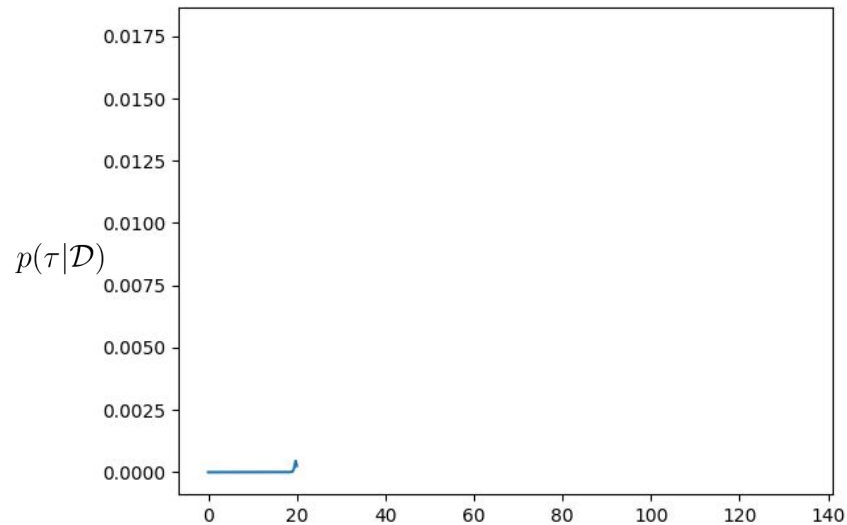
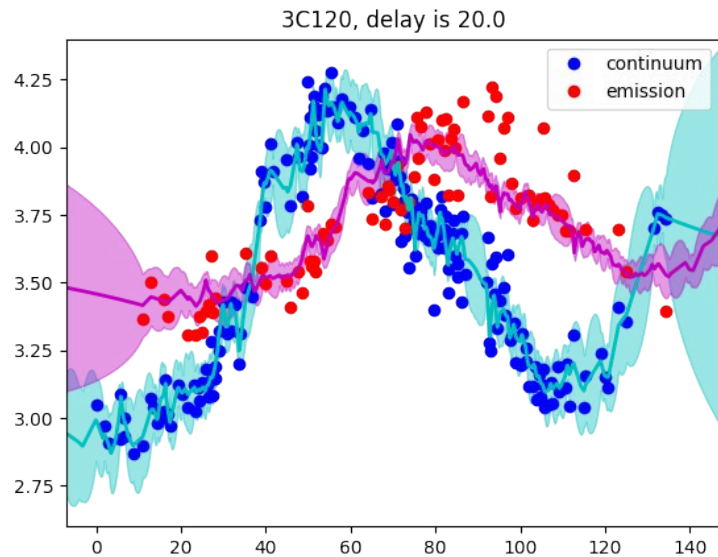




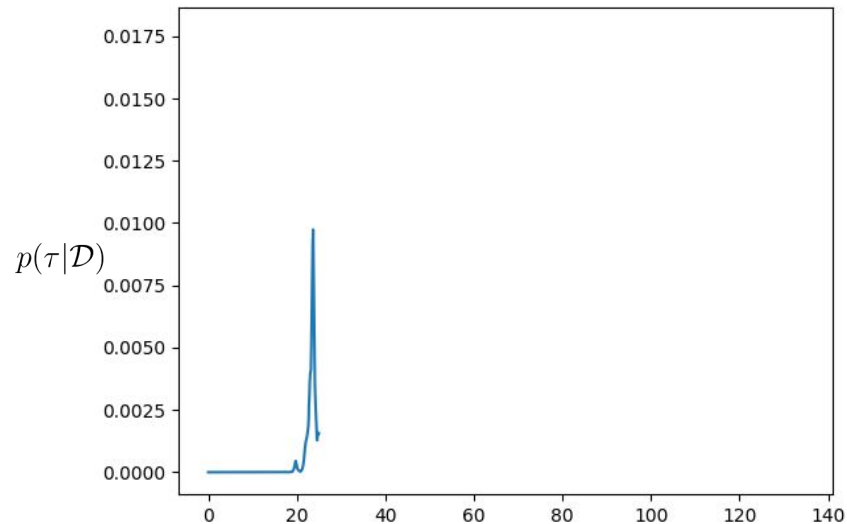
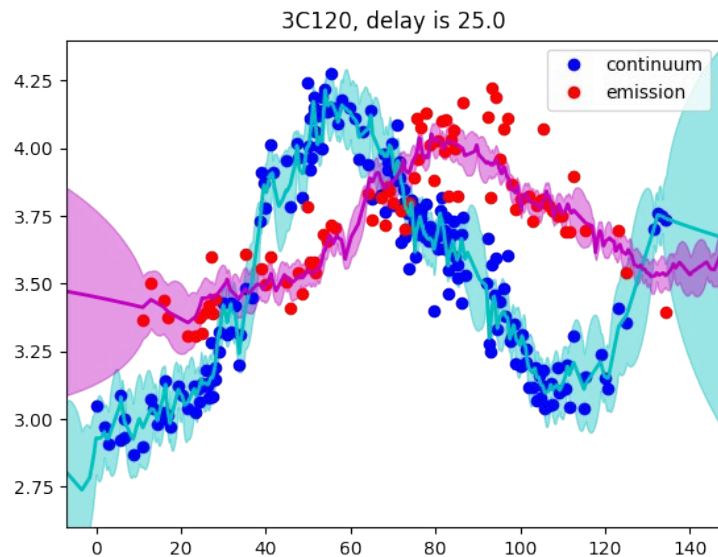
# Posterior delay distribution for 3C120



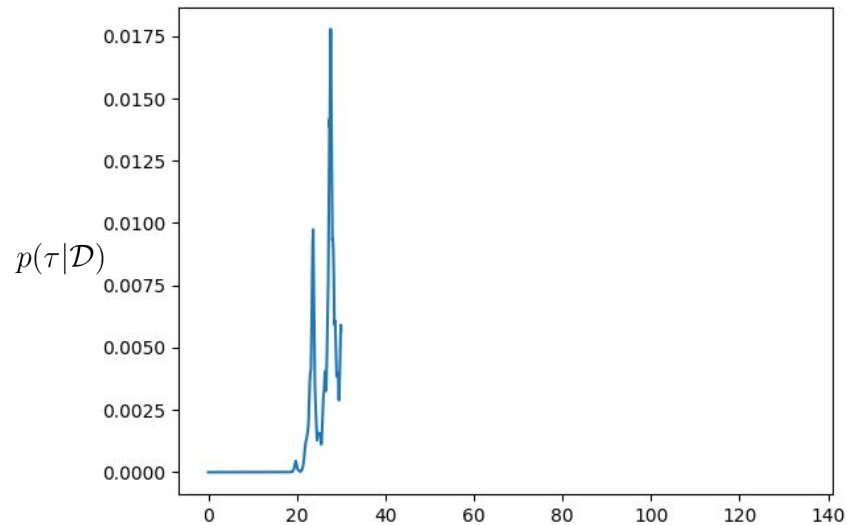
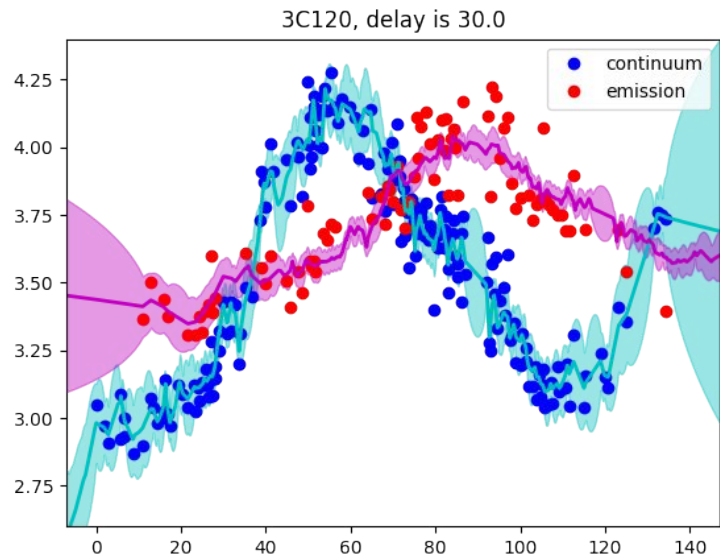
# Posterior delay distribution for 3C120



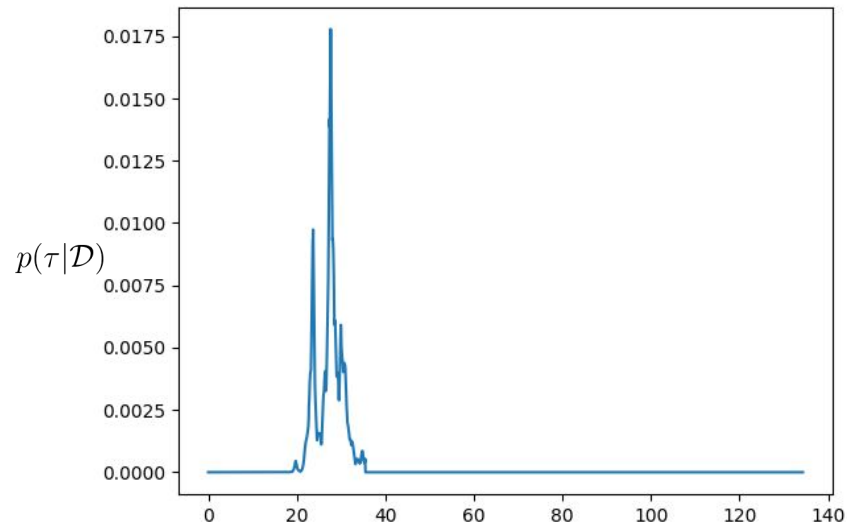
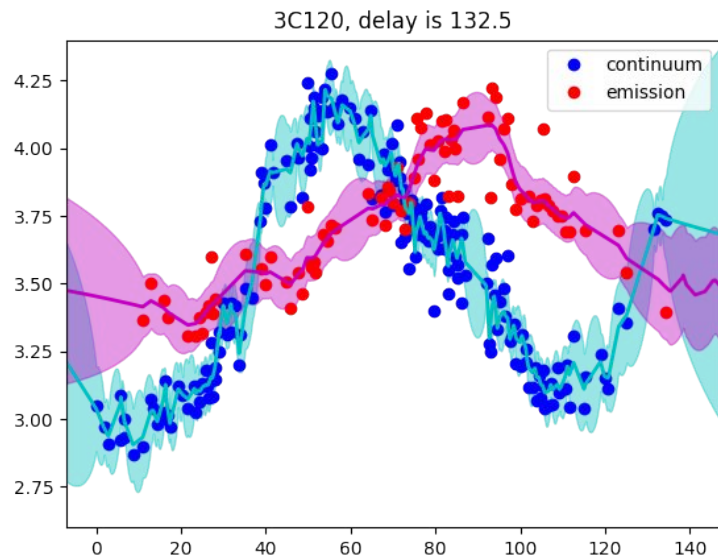
# Posterior delay distribution for 3C120



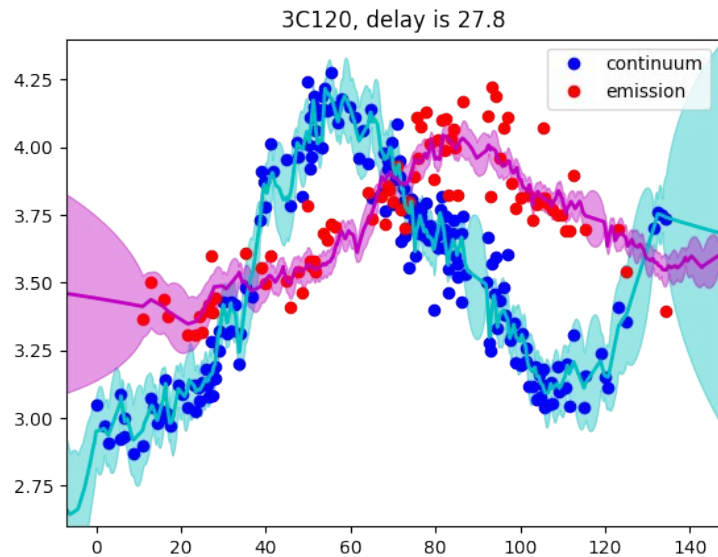
# Posterior delay distribution for 3C120



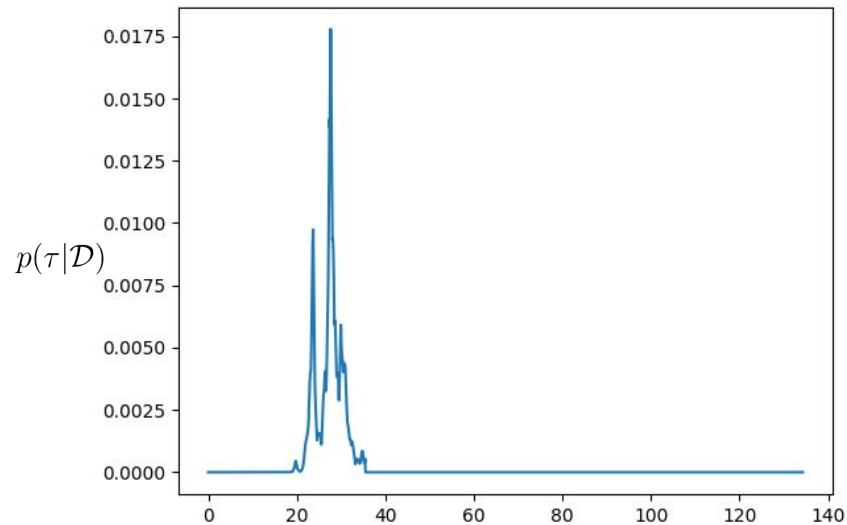
# Posterior delay distribution for 3C120



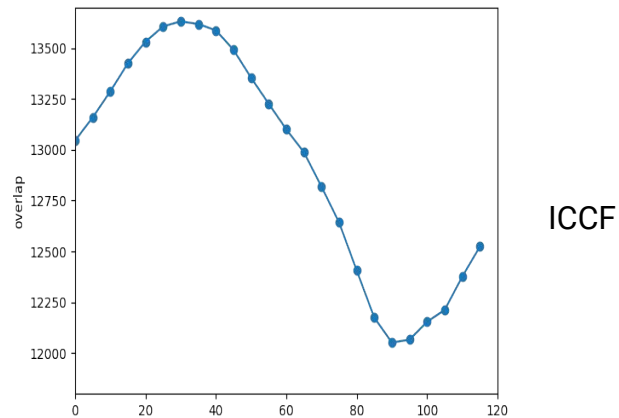
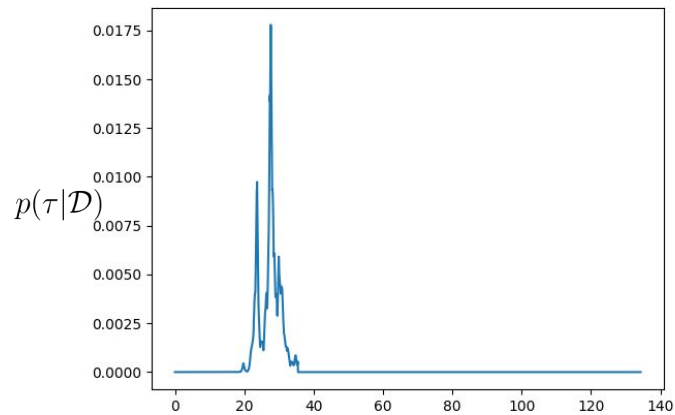
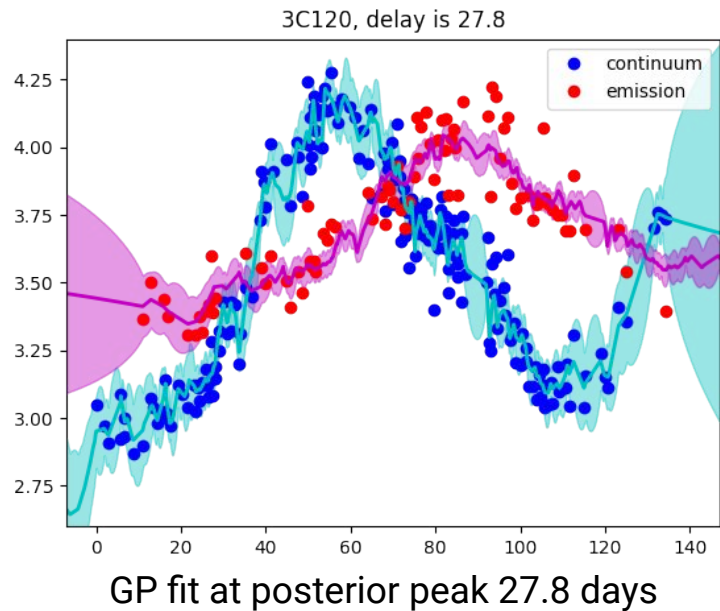
# Posterior delay distribution for 3C120



GP fit at posterior peak 27.8 days



# Posterior delay distribution for 3C120



# Summary

We reformulated ICCF in a probabilistic manner:

- Lightcurves related as  $y_e(t) = \alpha_e f(t - \tau) + b_e$   
 $y_c(t) = \alpha_c f(t) + b_c$
- Model  $f(t)$  as a GP
- Since GP closed under affine transformation, lightcurves also governed by GP
- We work out a posterior distribution for the delay

Our current model relies on simplification that one lightcurve is a **delayed**, **scaled** and **offsetted** version of the other.

In future work we want to be more physically realistic and use models of the type

$$y(t) = \int \mathcal{G}(\tau) f(t + \tau) d\tau$$

where a transfer function  $G$  encodes properties of the physical system